## Physics 371: Problem Set 7

Sean Carroll, Spring 2006 Due Thursday 1 June, 2:00 p.m.

1. (25 points) Consider a theory of two scalar fields,  $\phi_1$  and  $\phi_2$ , with potential

 $V(\phi_1, \phi_2) = (\phi_1^2 - \sigma^2)^2 + (\phi_2^2 - (\phi_1 - \sigma)^2)^2 .$ 

Write down the equations of motion for this theory. What are the possible zero-energy (vacuum) values of the fields? What kinds of domain walls are there in this theory?

- 2. (25 points) Imagine that cosmic strings are created by a phase transition at a temperature  $T \sim \sigma$ , with a tension  $\mu \sim \sigma^2$  (ignoring dimensionless parameters). But imagine that the interactions of the strings are such that they do *not* easily intercommute and chop up into loops, but instead get tangled up with each other, so that the overall string energy density redshifts as  $\rho_S \propto a^{-2}$ . What would the energy scale  $\sigma$  have to be for the strings to comprise the dark energy today (i.e.,  $\Omega_S = 0.7$ )? Approximately how likely is it that one such string is passing through the Solar System right now?
- 3. (50 points) Consider inflation with a potential  $V(\phi) = \frac{1}{2}m^2\phi^2$ . Imagine that inflation begins right at the Planck scale, with  $\rho \sim M_{pl}^4$ .
  - (a) Show that the field obeys  $\phi = \phi_0 \beta t$ , and solve for  $\beta$ .
  - (b) Show that the inflationary solution is stable to homogeneous perturbations. That is, let  $\phi(t) = \overline{\phi}(t) + \chi(t)$ , where  $\chi$  is a small perturbation. Derive an equation for  $\chi$ , and show that it doesn't grow with time.